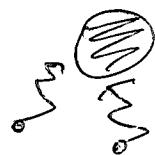


# Stochastic Processes

Example: Diffusion.



- Microscopic level

$$(L = 1\text{\AA} = 10^{-10}\text{m})$$

$\Rightarrow$  deterministic & chaotic

$$\dot{p} = -\frac{\partial H}{\partial q}$$

- Mesoscopic level

$$(L = 1\mu\text{m} = 10^{-6}\text{m})$$

$\Rightarrow$  stochastic

$$\delta \vec{x} = \delta \vec{F}$$

$\delta \vec{F}(t) =$  random force.

- symmetry: zero mean

$$\langle \delta \vec{F}(t) \rangle = 0$$

- noise strength:

$$\langle \int_{-\infty}^{\infty} \delta \vec{F}(t) \delta \vec{F}(t+\tau) d\tau \rangle = \langle |\tilde{\delta F}(0)|^2 \rangle = 2D.$$

- short-time correlations

$$\langle \delta \vec{F}(t) \delta \vec{F}(t+\tau) \rangle \approx 0 \text{ for } |\tau| \gg \sigma.$$

'back-of-the-envelope'

$$\sigma = \frac{L}{\sqrt{V}} = 1\text{nm},$$

$$V = \frac{6\pi r}{m}$$

Mean

$$\langle x(t) \rangle = 0.$$

Mean-square displacement:

$$\langle x(t)^2 \rangle =$$

$$\langle \left( \int_0^t dt_1 \dot{x}(t_1) \right) \left( \int_0^t dt_2 \dot{x}(t_2) \right) \rangle =$$

$$= \int_0^t dt_1 \int_0^t dt_2 \langle \dot{x}(t_1) \dot{x}(t_2) \rangle =$$

$$= \int_0^t dt_1 \int_{-t_1}^{t-t_1} dt_2 \langle S(t_1) S(t_1+t_2) \rangle$$

$$\approx \int_0^t dt_1 \int_{-\infty}^{\infty} dt_2 \langle S(t_1) S(t_1+T) \rangle.$$

$$= 2D t.$$

$D$  = diffusion coefficient

What is D?

Great Idea (Einstein 1905)

Use equipartition theorem.

Trick to exploit idea:

add elastic spring.



$$\text{Stokes } \gamma = 6\pi\eta a$$

Oscillated motion.

$$(*) \quad \ddot{x} = -\frac{k}{\gamma}x + \xi(t) \equiv \text{Ornstein-Uhlenbeck process}$$



$$\langle x(t) \rangle = 0$$

Equipartition theorem:

$$\frac{k}{2} \langle x^2 \rangle = \frac{\xi_0^2}{2}$$

Final solution of (\*).

$$x(t) = \int_{-\infty}^t \xi(\tau) \exp -\frac{k(t-\tau)}{\gamma} d\tau$$

Now

$$\langle X(t)^2 \rangle =$$

$$= \left\langle \left( \int_{-\infty}^t dt_1 S(t_1) \exp - \frac{\zeta(t-t_1)}{\gamma} \right) \right. \\ \left. \left( \int_{-\infty}^t dt_2 S(t_2) \exp - \frac{\zeta(t-t_2)}{\gamma} \right) \right\rangle$$

$$= \int_{-\infty}^t dt_1 \exp - \frac{2\zeta(t-t_1)}{\gamma}$$
$$\underbrace{\int_{-\infty}^{t-t_1} d\tau \langle S(t_1) S(t_1+\tau) \rangle \exp + \frac{\zeta \tau}{\gamma}}_{\approx 2D} \quad (\text{We need } \delta \ll \gamma/k)$$

$$\approx 2D \cdot \left[ \frac{\gamma}{2k} \exp - \frac{2\zeta(t-t_1)}{\gamma} \right]_t^{-\infty}$$

$$= 2D \cdot \frac{\gamma}{2k}$$

$$\frac{k}{2} \cdot \langle X^2 \rangle = \frac{k}{2} \cdot 2D \cdot \frac{\gamma}{2k} = \frac{k_B T}{2}$$
$$\Rightarrow \text{Einstein relation}$$

$$\boxed{D = \frac{k_B T}{\gamma}}$$

Special instance of  $\overline{D}$ .

+ Fluct. spectrum = Dissipation spectrum

What is the structure of the noise term  $\xi(t)$ ?

### Mathematical view

$$W(t) = \int_0^t d\tau \xi(\tau).$$

$$\langle W(t) \rangle = 0.$$

$$\langle W(t)^2 \rangle = \int_0^t dt_1 \int_0^{t_1} dt_2 \langle W(t_1) W(t_2) \rangle$$

$$= \int_0^t dt_1 \int_{t_1}^{t-t_1} dt \langle W(t_1) W(t+\tau) \rangle$$

$$\approx \int_0^t dt_1 \int_{-\infty}^{\infty} d\tau \langle W(t_1) W(t_1+\tau) \rangle$$

$$= 2D t.$$

What about  $\langle W(t)^2 \rangle = ?$

⇒ realization:

$$dW = W(t+dt) - W(t)$$

≡ Gaussian random variable  
with mean zero

and variance  $2D dt$ .

≡ Wiener process.

## Physicist's View

$\xi(t) = \text{Gaussian white noise}$   
 $\sigma \rightarrow 0, \Rightarrow$   
δ-correlated noise.

$$\langle \xi(t) \rangle = 0.$$

$$\langle \xi(t_1) \xi(t_2) \rangle = 2D \delta(t_1 - t_2).$$

$$\langle \xi(t_1) \cdot \xi(t_2) \cdot \dots \cdot \xi(t_m) \rangle$$

$$= \begin{cases} 0 & m = 2n+1 \\ \sum_i (2D)^n \delta(t_{i_1} - t_{i_2}) \dots \delta(t_{i_{m-1}} - t_{i_m}) & m = 2n \end{cases}$$

## Numerics

Example:  $\dot{X} = f(t)$ ,

$$x_n = x(t_n), \quad t_n = n dt.$$

Euler scheme:

$$x_{n+1} = x_n + \sqrt{2D dt} W_n.$$

With  $W_n$  normal distributed random variable with mean zero and variance  $1$ .

N.B.  $x_{n+1} - x_n = \sqrt{2D} \cdot dW$ .

$$\langle dW^2 \rangle = 2D dt.$$

General case:

$$\dot{X} = f(x) + I(t).$$

$$x_{n+1} = x_n + f(x_n) dt + \sqrt{2D dt} W_n.$$

Caution: Take care of  $D=D(x)$ .

$P(x,t)$  = probability density.

Example:  $\dot{x} = \xi(t)$ ,  $x(0) = 0$ .

$$\Rightarrow P(x,t) = \frac{1}{\sqrt{2\pi} 2D_t} \exp\left(-\frac{x^2}{4D_t}\right)$$

$$\frac{\partial}{\partial t} P(x,t) = D \frac{\partial^2}{\partial x^2} P(x,t).$$

Check:

$$\begin{aligned} \frac{\partial}{\partial t} P + \frac{1}{4D} \frac{x^2}{t^2} P, &= \frac{\partial}{\partial x} P = \frac{-2x}{4Dt} P, \\ \frac{\partial^2}{\partial x^2} P &= \frac{-2}{4Dt} P + \left(\frac{-2x}{4Dt}\right)^2 P. \end{aligned}$$

General Case:

$$\dot{x} = f(x) + \xi(t).$$

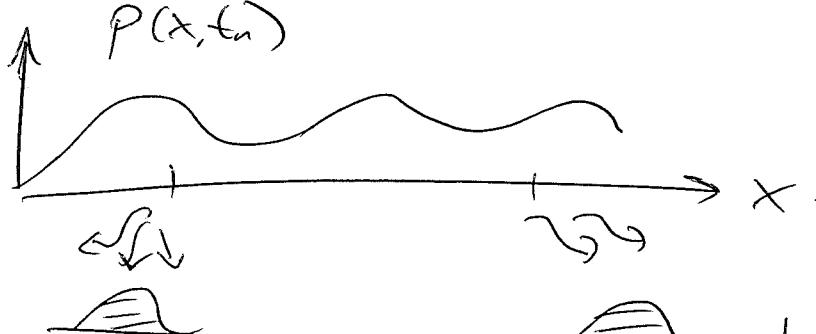
$$\frac{\partial}{\partial t} P(x,t) = L \cdot P(x,t).$$

$$L = -\frac{\partial}{\partial x} f + \frac{\partial}{\partial x} D \frac{\partial}{\partial x}.$$

Caution: if  $D=D(x)$ !

Q: How to find  $L$ ?

A: Propagate  $P(x, t_n)$  to  $P(x, t_{n+1})$ .



$$N\left(\frac{x - x_n - f(x_n)dt}{\sqrt{2Ddt}}\right)$$

= normal distribution

$$x_{n+1} = x_n + f(x_n)dt + \sqrt{2Ddt} N.$$

Makov property  $\Rightarrow$

Dogmen-Volenskar equation

$$P(x, t_{n+1}) = \int dx_n P(x_n, t_n) \cdot N\left(\frac{x - x_n - f(x_n)dt}{\sqrt{2Ddt}}\right)$$

probability to  
have been at  
 $x_n$  at  $t=t_n$

probability to have  
moved from  
 $x_n$  to  $x_{n+1}$ .

We restrict ourselves to case

$$D(x) = D_0.$$

$$\text{Let } y = x - x_n.$$

$$p(y) = P(x-y, t_n)$$

$$n(y) = N\left(\frac{z - f(x-y)dt}{\sqrt{2Ddt}}\right)$$

$$\text{integrand} = p(y) \cdot n(y)$$

$$= P \cdot n|_{y=0}$$

$$+ \frac{\partial}{\partial x} (P \cdot n)|_{y=0} \cdot y$$

$$+ \frac{\partial^2}{\partial x^2} (P \cdot n)|_{y=0} \cdot \frac{y^2}{2} + \dots$$

$$P(x, t_{n+1}) = \int dx_n \quad \text{integrand} \quad \text{if we let } z=y.$$

$$= \int dy \quad \text{integrand}$$

$$= P \cdot \int dy \cdot n$$

$$+ \frac{\partial}{\partial x} [P \int dy \cdot n \cdot y]$$

$$+ \frac{\partial^2}{\partial x^2} [P \int dy \cdot n \cdot \frac{y^2}{2}] + \dots$$

$$\left[ \int dy \cdot n = 1, \quad \int dy \cdot n \cdot y = f(x)dt, \quad \int dy \cdot n \cdot y^2 = 2Ddt + [f(x)dt]^2 \right]$$

$$\dots = P + \frac{\partial}{\partial x} [P \cdot f(x)] dt + \frac{\partial^2}{\partial x^2} P D dt$$

$$\frac{P(x, t_{n+1}) - P(x_n, t_n)}{dt} = -\frac{\partial}{\partial x} (P \cdot f(x)) + D \frac{\partial^2}{\partial x^2} P(x, t).$$

## Application

Diffusion in potential  $U(x)$ .

$$\dot{X} = -\frac{1}{\gamma} \frac{\partial U}{\partial x} + \xi.$$

Steady state:  $\frac{\partial}{\partial x} P = 0$ ,

$$0 = + \nabla \left[ \left( \frac{1}{\gamma} \nabla u \right) P \right] + D \nabla^2 P.$$

$$= \nabla \left[ \left( \frac{1}{\gamma} \nabla u \right) P + D \nabla P \right].$$

$$T(x) \equiv T_0 \Rightarrow D = \frac{k_B T_0}{\gamma}$$

$\Rightarrow$  Boltzmann distribution.

$$\text{Check: } \frac{P \sim \exp - \frac{U(x)}{k_B T}}{1}$$

$$\nabla P \sim -\frac{1}{k_B T} P \cdot (\nabla u).$$

$$\Rightarrow \left( \frac{1}{\gamma} \nabla u \right) P + D \nabla P = 0.$$

Fokker-Planck equation = .

Conservation equation

$$\dot{P} = -\nabla \mathcal{F}.$$

Different b.c. - different phys. meaning.

• Reflecting b.c.

$$\mathcal{F}(x=0) = \mathcal{F}(x=L) = 0.$$

$$\Rightarrow \frac{d}{dt} \int_0^L dx P = 0.$$

$$\dot{P} = L P, \quad \lambda P = L P \Rightarrow \lambda_1 \geq \lambda_2 \geq \dots$$

• Absorbing b.c.  $\Rightarrow \lambda_1 = 0, P_1 = P^*$   
 $\lambda_i < 0$  for  $i > 1$ .

$$P(x=L) = 0.$$

$$\frac{d}{dt} \int_0^L dx P(x,t) = -\mathcal{F}_L(t)$$

$\equiv$  flux to absor.

$\lambda < 0$   $\equiv$  lowest time-scale.

$\leftarrow \frac{1}{\lambda_1}: P(x,t) \rightarrow p_1(t),$

$$\dot{p}_1(t) \sim \exp(-\lambda_1 t).$$